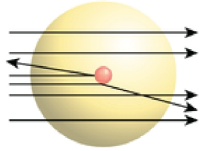


RUTHERFORD'S ATOM MODEL

- i) Majority of α - particles passed without any deviation.
- ii) Some are scattered at small angle θ (impact parameter is equal to that of nuclear radius)
- iii) Only few alpha particle retrace the path (impact parameter = 0)

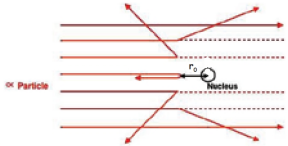


DISTANCE OF CLOSEST APPROACH OF α -PARTICLES

$$r_0 = \frac{1}{4\pi\epsilon_0} \frac{4Ze^2}{mv^2}$$

$$r_0 \propto \frac{1}{m} \quad r_0 \propto \frac{1}{v^2}$$

$$r_0 \propto \frac{1}{K.E} \quad r_0 \propto Z_1 Z_2$$

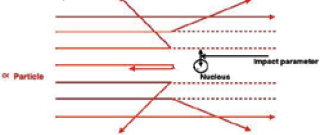


IMPACT PARAMETER

$$b = \frac{1}{4\pi\epsilon_0} \frac{Ze^2 \cot \frac{\theta}{2}}{\frac{1}{2} mv^2}$$

$$b \propto \frac{1}{m} \quad b \propto \frac{1}{v^2}$$

$$b \propto \frac{1}{K.E} \quad b \propto \cot \frac{\theta}{2}$$



BOHR ATOM MODEL

First postulate

$$F = \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r^2}$$

Second postulate

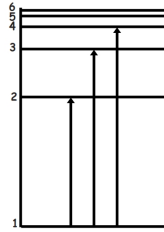
$$mvr = \frac{nh}{2\pi}$$

$$\frac{mv^2}{r} = \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r^2}$$

ATOMS

HYDROGEN SPECTRUM

Absorption spectrum



Electrons absorb only those photons whose energy = Energy difference of 2 shells

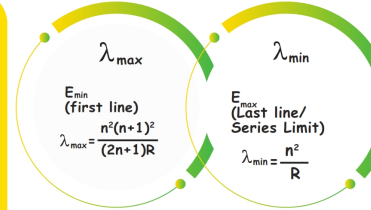
If atomic excitation takes place upto n^{th} shell starting from ground state then $(n-1)$ different photons are absorbed

ENERGY

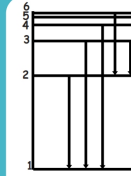
$$\text{Total energy} = -13.6 \frac{Z^2}{n^2} \text{ eV}$$

$$K.E = -T.E = +13.6 \frac{Z^2}{n^2} \text{ eV}$$

$$P.E = 2T.E$$



EMISSION SPECTRUM



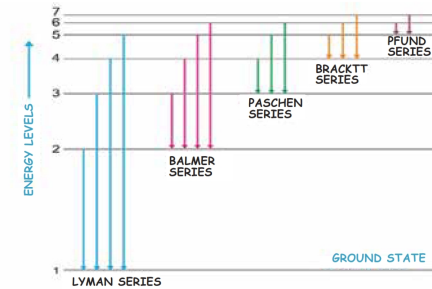
Total no. of different wavelength photons in emission spectrum = $\frac{n(n-1)}{2}$

All absorbed photons are emitted in emission spectrum

Wavelength of emitted photon

$$\bar{\nu} = \frac{1}{\lambda} = R Z^2 \left[\frac{1}{n_1^2} - \frac{1}{n_2^2} \right] \quad R \approx 10^5 \text{ cm}^{-1} \\ \frac{1}{R} = 91 \text{ nm} = 910 \text{ \AA} \quad R \approx 10^7 \text{ m}^{-1}$$

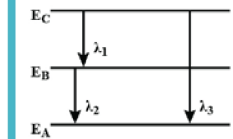
LINE SPECTRUM OF HYDROGEN ATOM



LINE SPECTRUM OF HYDROGEN ATOM

Spectral series	n_1	n_2	Wavelength	$\lambda_{\text{max}} (n_2 = n_1 + 1)$	$\lambda_{\text{min}} (n_2 = \infty)$	Region	Range
Lyman	1	2, 3, 4	$\frac{1}{\lambda_{\text{Ly}}} = R \left(\frac{1}{1^2} - \frac{1}{n^2} \right)$	$\frac{4}{3R}$	$\frac{1}{R}$	Ultra - violet	911.6 \AA to 1216 \AA
Balmer	2	3, 4, 5	$\frac{1}{\lambda_{\text{B}}} = R \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$	$\frac{36}{5R}$	$\frac{4}{R}$	Visible	3646 \AA to 6563 \AA
Paschen	3	4, 5, 6	$\frac{1}{\lambda_{\text{P}}} = R \left(\frac{1}{3^2} - \frac{1}{n^2} \right)$	$\frac{144}{7R}$	$\frac{9}{R}$	Near infra-red	8204 \AA to 18753 \AA
Brackett	4	5, 6, 7	$\frac{1}{\lambda_{\text{Br}}} = R \left(\frac{1}{4^2} - \frac{1}{n^2} \right)$	$\frac{400}{9R}$	$\frac{16}{R}$	Middle infra-red	14585 \AA to 40515 \AA
Pfund	5	6, 7, 8	$\frac{1}{\lambda_{\text{Pf}}} = R \left(\frac{1}{5^2} - \frac{1}{n^2} \right)$	$\frac{900}{11R}$	$\frac{25}{R}$	Far infra-red	22790 \AA to 74583 \AA

Energy levels A, B & C of a certain atom correspond to increasing values of energy, i.e. $E_A < E_B < E_C$. If $\lambda_1, \lambda_2, \lambda_3$ are the wavelengths of radiations corresponding to transitions C to B, B to A and C to A respectively then



a) $\lambda_3 = \lambda_1 + \lambda_2$ c) $\lambda_1 + \lambda_2 + \lambda_3 = 0$
 b) $\lambda_3 = \frac{\lambda_1 \lambda_2}{\lambda_1 + \lambda_2}$ d) $\lambda_3^2 = \lambda_1^2 + \lambda_2^2$